

An Introduction to symptotic pproximants

Nate Barlow
School of Mathematical Sciences, RIT

All work done with co-conspirator of the method:
Steve Weinstein
Chemical Engineering, RIT

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Overview

In this talk, the method of *asymptotic approximants* will be applied to obtain analytic solutions for:

- ▶ an elliptic **integral** important to light bending.
- ▶ nonlinear **differential equations** of fluid dynamics.
- ▶ the analytic continuation of certain **truncated expansions** of thermodynamics and astrophysics

History of Asymptotic approximants

- ▶ 1961: Baker & Gammel's Padé Approximant theorem states when an approximant is expected to converge.
⋮
- ▶ Jumps and Starts that invoke Baker & Gammel's idea, but nothing formalized as a method
⋮
- ▶ 2012: (Barlow et al, J. Chem. Phys., 2012) An asymptotically consistent approximant for soft-sphere fluids
- ▶ 2013-2015: Additional papers in thermodynamics start to form a unified approach to solving problems where behavior is known at two ends.
- ▶ 2015-2017: Method, now coined "asymptotic approximants", applied to boundary layer problems in fluid dynamics. Leads to a methods Paper (Barlow et al, QJMAM, 2017).
- ▶ 2017-present: Nate and Steve start working with Josh as the method is applied to astrophysics. Leads to two papers on light bending and a collaboration with Ofek & CCRG folks on application to gravitational waves.

Example: Integral for light bending

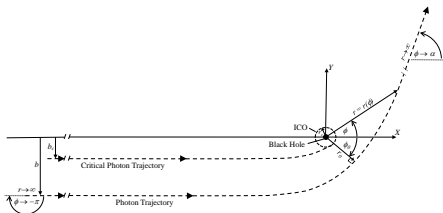
The governing equation for trajectories of light around a Kerr black-hole in the equatorial plane is given by

$$\phi(r) = \ll \pi + \int_0^{r_0/r} g(\hat{y}) d\hat{y}$$

where

$$g(\hat{y}) = \frac{u_0 [b \ll 2u_0 (b \ll a) \hat{y}]}{[1 \ll 2u_0 \hat{y} + a^2 u_0^2 \hat{y}^2] \Sigma + \hat{y} \ll 2(b \ll a)^2 u_0^3 \hat{y}^2}^{1/2} (1 \ll \hat{y})^{1/2},$$

$u_0 = 1/r_0$, a is spin, and other parameters (b , r_0) defined below.

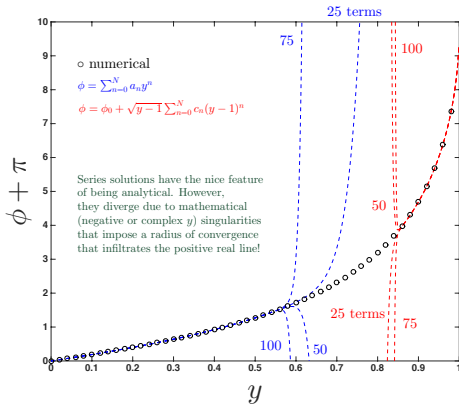


Example: Integral for light bending

Let $y \equiv r_0/r$.

$$\phi(y) = \llcorner\pi + \int_0^y g(\hat{y})d\hat{y}$$

What can we do to solve this?



Solution shown here for $a = 1$, $b = 2.2222$.

Example: Integral for light bending

We can use the analytical information from the series, get past the radius of convergence, and bridge the gap in the solution as follows:

- ▶ We know that the expansion of the integral about $y = 0$ is

$$\phi = \sum_{n=0}^{\infty} a_n y^n$$

- ▶ We also know that the leading order behavior about $y = 1$ is

$$\phi = \phi_0 + c_0 \sqrt{y} \ll 1$$

- ▶ Lets make up a function that (when expanded about $y = 0$) limits to the series above while also limiting to the square root behavior as $y \rightarrow 1$. How about this (bear with me here):

$$\phi_A = \phi_0 + \sqrt{1 \ll y} \sum_{n=0}^{\infty} A_n (y \ll 1)^n$$

- ▶ In this, we have already satisfied the correct leading $y \rightarrow 1$ behavior.
- ▶ The A_n are not known at this point! We will choose them in order to meet our other goal ...

Example: Integral for light bending

$$\phi_A = \phi_0 + \sqrt{1 \ll y} \sum_{n=0}^{\infty} A_n (y \ll 1)^n$$

We “choose” the A_n 's such that the expansion of the above about $y = 0$ equals the expansion of the original integral about $y = 0$, namely:

$$\phi = \sum_{n=0}^{\infty} a_n y^n.$$

This can be viewed as solving a linear system (replace ∞ with N)

$$\phi_A(0) = f_0(A_0 \dots A_N) = a_0$$

$$\phi'_A(0) = f_1(A_0 \dots A_N) = a_1$$

$$\phi''_A(0) = f_2(A_0 \dots A_N) = 2 a_2$$

\vdots

$$\phi_A^{(N)}(0) = f_N(A_0 \dots A_N) = N! a_N$$

But a matrix inversion is not saving us much (if any) over numerical integration of the original integral!! We can do better ...

Example: Integral for light bending

$$\phi_A = \phi_0 + \sqrt{1-y} \sum_{n=0}^{\infty} A_n (y-1)^n$$

1. Replace the LHS of above with expansion about $y = 0$

$$\sum_{n=0}^{\infty} a_n y^n = \phi_0 + \sqrt{1-y} \sum_{n=0}^{\infty} A_n (y-1)^n$$

2. Isolate the A_n series on one side

$$(1-y)^{-1/2} \left[-\phi_0 + \sum_{n=0}^{\infty} a_n y^n \right] = \sum_{n=0}^{\infty} A_n (y-1)^n$$

3. Expand the LHS of above about $y = 0$

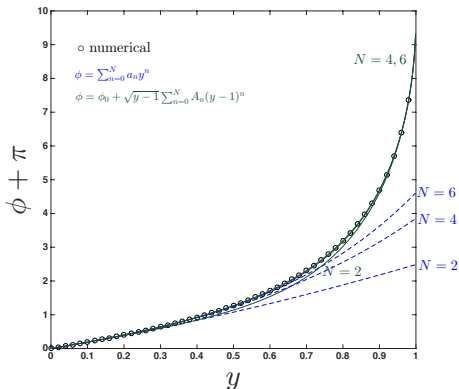
$$\sum_{n=0}^{\infty} G_n y^n = \sum_{n=0}^{\infty} A_n (y-1)^n$$

4. Replace ∞ with N and this becomes an identity with the solution for the coefficients

$$A_n = \frac{1}{n!} \sum_{m=0}^N \frac{\Gamma(m+1)}{\Gamma(m-n+1)} G_m$$

Example: Integral for light bending

The *Approximant* doesn't take that many terms before locking in!



Solution shown here for $a = 1$, $b = 2.2222$.

(show trajectory animation)

More details in:

R. J. Beachley, M. Mistysyn, J. A. Faber, S. J. Weinstein, and N. S. Barlow. Accurate closed-form trajectories of light around a kerr black hole using asymptotic approximants. *Class. Quant. Grav.*, 35(20):128, 2018.

Obligatory Sherlock Holmes Quote

"See the value of imagination... We imagined what might have happened, acted upon the supposition, and found ourselves justified"

- A. C. Doyle from *"Silver Blaze"* (1892)

Asymptotic Approximants

- ▶ Hybrid of two well-known mathematical techniques
 - ▶ Asymptotic Matching - combining two *overlapping* asymptotic expansions
 - ▶ Padé Approximants - series expansion of rational function (polynomial over polynomial) about given point is same as true expansion about that point
- ▶ May be constructed when asymptotic behaviors are known in two different regions of a domain.
- ▶ Definition:
 - ▶ Form of approximant matches the behavior in one limit
 - ▶ Unknowns are chosen to match the behavior in the other limit

Definition (Barlow et al, QJMAM 2017)

Given a power series representation of some function $f(x)$:

$$f = \sum_{n=0}^{\infty} a_n(x - x_0)^n, \quad (1)$$

and an asymptotic behavior

$$f \rightarrow f_a(x) \text{ as } x \rightarrow x_a,$$

where $x_a \neq x_0$, an *asymptotic approximant* is any function $f_A(x)$ that may be expressed analytically in closed form and that satisfies the following three properties:

1. The N -term Taylor expansion of f_A about x_0 is identical to the N -term truncation of (1).
2. $\lim_{x \rightarrow x_a} (f_A/f_a) = \text{constant}$ for any N .
3. The sequence of approximants converges for increasing N .

Asymptotic approximants approximate and constrain the analytic continuation of an expansion such that the correct asymptotic limit (in some other region) is obtained.

Flow near a flat plate

steady, incompressible, 2-D, neglect gravity, $v \ll u$, $\frac{\partial}{\partial x} \ll \frac{\partial}{\partial y}$

governing “boundary layer” equations:

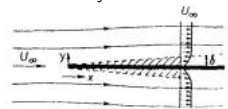
$$\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0$$

$$u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} = \nu \frac{\partial^2 u}{\partial y^2}$$

boundary conditions:

$$u = v = 0 \text{ at } y = 0$$

$$u \rightarrow U \text{ as } y \rightarrow \infty$$



Blasius (1908)

picture from Schlichting

$$u = U, v = 0 \text{ at } y = 0$$

$$u \rightarrow 0 \text{ as } y \rightarrow \infty$$

or



Sakiadis (1961)

Similarity transform

$$\eta = y\sqrt{U/(\nu x)}, \quad u = Uf'(\eta), \quad v = [\eta f'(\eta) - f(\eta)]\sqrt{\nu U/(4x)}$$

single nonlinear ordinary differential equation in $f(\eta)$:

$$2f''' + ff'' = 0$$

boundary conditions:

Blasius Problem



$$f(0) = f'(0) = 0, \quad f'(\infty) \rightarrow 1$$

Sakiadis Problem



$$f(0) = 0, \quad f'(0) = 1, \quad f'(\infty) \rightarrow 0$$

Power Series Solution to the Blasius/Sakiadis Problem

$$2f''' + ff'' = 0$$

$$f(0) = a_0, f'(0) = a_1, f'(\infty) \rightarrow b$$

$$f = \sum_{n=0}^{\infty} a_n x^n$$

$$a_{n+3} = \frac{-\sum_{j=0}^n a_{n-j}(j+1)(j+2)a_{j+2}}{2(n+1)(n+2)(n+3)}.$$

- ▶ requires $f''(0) \equiv \kappa$, coefficient of the wall shear

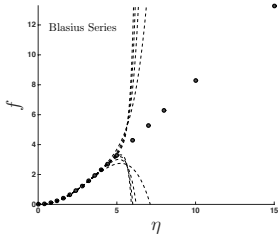
$$f = a_0 + a_1\eta + \frac{\kappa}{2}\eta^2 + a_3(\kappa)\eta^3 + a_4(\kappa)\eta^4 + \dots$$

- ▶ numerical estimates of κ :
 - ▶ Blasius flow: 0.33205733621519630 (Boyd, 1999)
 - ▶ Sakiadis flow: -0.44374733 (Cortell, 2010)

Power Series Solution to the Blasius/Sakiadis Problem

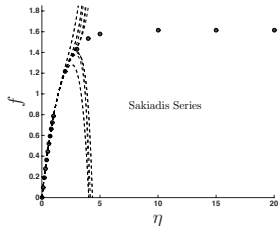
$$2f''' + ff'' = 0$$

Blasius Problem



$$f(0) = f'(0) = 0, f'(\infty) \rightarrow 1$$

Sakiadis Problem



$$f(0) = 0, f'(0) = 1, f'(\infty) \rightarrow 0$$

Approximant for the Blasius Problem

We start with an approximant form that matches $f'(\infty) \rightarrow 1$:

$$f_A = \eta + B \ll B \left(1 + \sum_{n=1}^N A_n \eta^n \right)^{-1}$$

and follow steps from before to find the A_n 's:

1.

$$\sum_{n=0}^{\infty} a_n \eta^n = \eta + B \ll B \left(1 + \sum_{n=1}^N A_n \eta^n \right)^{-1}$$

2.

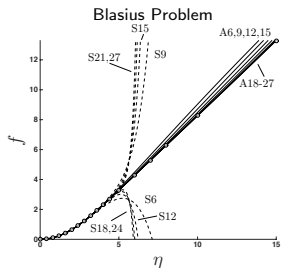
$$\left[\eta/B + 1 \ll \frac{1}{B} \sum_{n=0}^{\infty} a_n \eta^n \right]^{-1} = 1 + \sum_{n=1}^N A_n \eta^n$$

3. Expand LHS about $\eta=0$. JCP Miller's formula leads to a recursive solution:

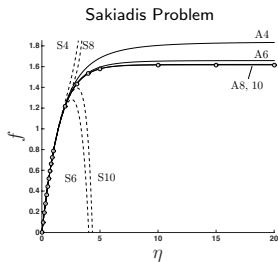
$$A_{n>0} = \frac{1}{B} \sum_{j=1}^n \tilde{a}_j A_{n-j}, \quad A_0 = 1, \quad \tilde{a}_1 = \ll 1, \quad \tilde{a}_{j>1} = a_j$$

Approximants to the Blasius/Sakiadis Problem

$$2f''' + ff'' = 0$$



$$f(0) = f'(0) = 0, f'(\infty) \rightarrow 1$$

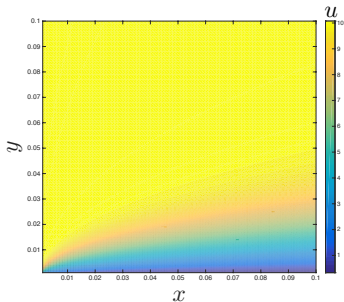


$$f(0) = 0, f'(0) = 1, f'(\infty) \rightarrow 0$$

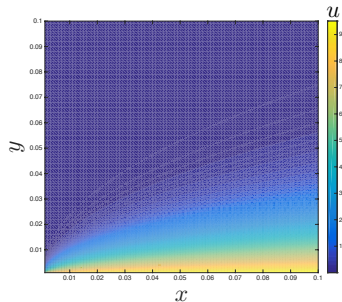
(Barlow et al, QJMAM 2017)

Transform back to physical variables

Blasius Problem



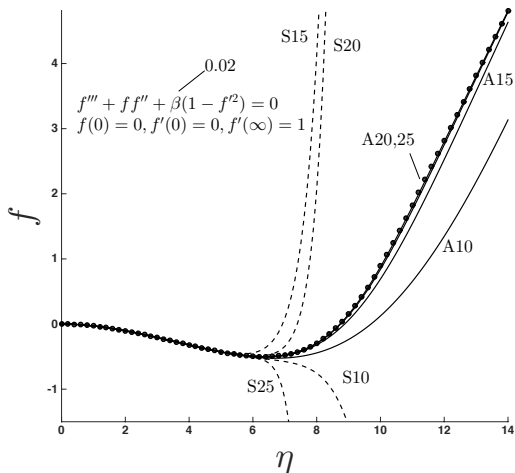
Sakiadis Problem



(Barlow et al, QJMAM 2017)

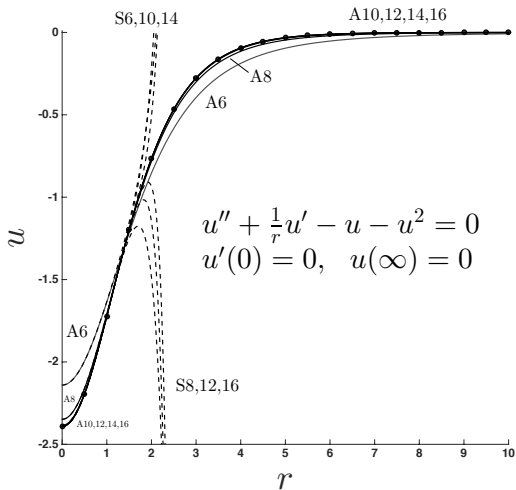
Some other nonlinear ODEs we've solved:

Falkner-Skan equation for boundary layers over a wedge



Some other nonlinear ODEs we've solved:

Flieril-Petviashvili model for the motion of Jupiters red spot

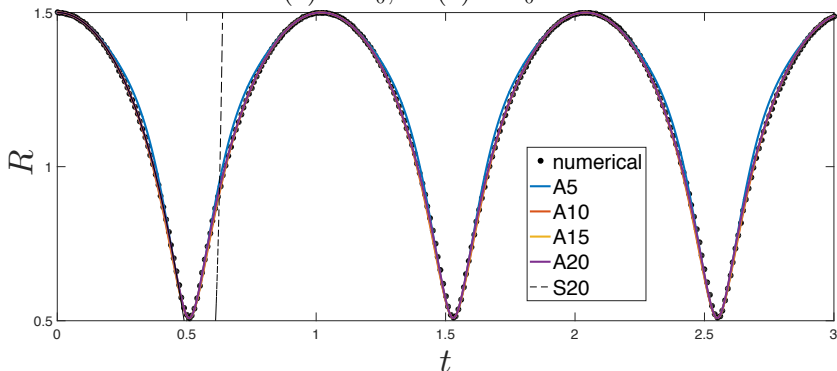


Some other nonlinear ODEs we've solved:

Currently Working on w/ A. Harkin, A. Giammarese, J. Tarantino:

Rayleigh-Plesset Equation for Oscillating Bubbles

$$R\ddot{R} + \frac{3}{2}(\dot{R})^2 + \delta\frac{1}{R} - \alpha\beta\frac{1}{R^3} = \gamma$$
$$R(0) = R_0, \quad R'(0) = V_0$$



Let's try some harder problems. . .

What if you have no governing equation to start from (integral, differential equation, etc.) ?

What if you only have a **limited number of terms in an expansion**. . . and the expansion diverges?! (at some point)

This is the case for the **virial expansion** equation of state from thermodynamics.

$$\frac{P}{\rho kT} = 1 + B_2\rho + B_3\rho^2 + \dots$$

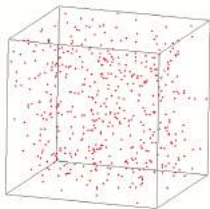
calculating fluid properties

Properties such as energy, entropy, etc. can be calculated by applying thermodynamic laws to an equation of state.

- ▶ ex. ideal gas

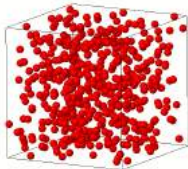
$$P = \rho kT$$

- ▶ assumes no interaction b/w fluid molecules
- ▶ valid when molecules are “far enough” apart (limit as $\rho \rightarrow 0$)



virial series

What if molecular interactions are important? (non-ideal fluid)



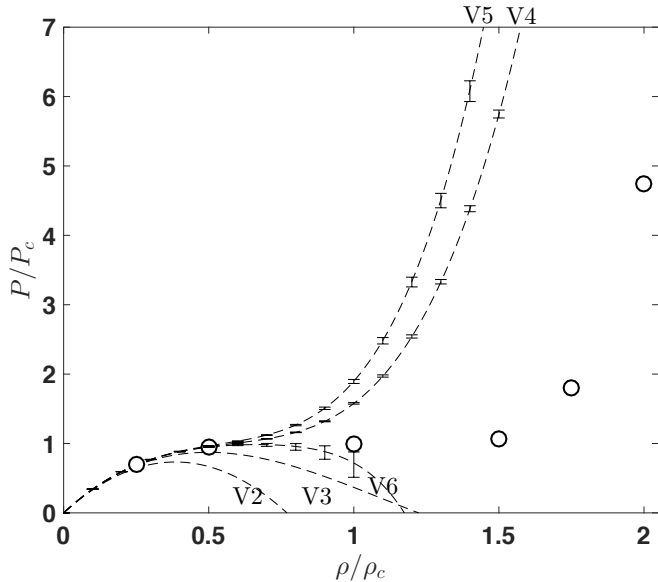
- ▶ virial equation of state (virial series)

$$P = kT \sum_{j=1}^J B_j(T) \rho^j$$

$$B_1 = 1, \quad B_2 = -\frac{1}{2} \int \tilde{f}_{1,2} d\mathbf{r}_{1,2}, \dots \quad B_8 = 18\text{-dimensional integral}$$

- ▶ intermolecular interactions accounted for
- ▶ expansion of $P(\rho, T)$ about $\rho=0$ (ideal gas limit)

virial series for a model fluid (square-well)



critical point & scaling

- ▶ The thermodynamic critical point $(\rho, T, P) = (\rho_c, T_c, P_c)$ occurs where

$$\left(\frac{\partial P}{\partial \rho}\right)_{T,N} = 0; \quad \left(\frac{\partial^2 P}{\partial \rho^2}\right)_{T,N} = 0; \quad \left(\frac{\partial^3 P}{\partial \rho^3}\right)_{T,N} \geq 0.$$

- ▶ universal critical scaling:

$$(P - P_c)_{T_c} \sim C \operatorname{sgn}\left(1 - \frac{\rho}{\rho_c}\right) \left|1 - \frac{\rho}{\rho_c}\right|^\delta \quad \text{as } \rho \rightarrow \rho_c$$

- ▶ (Pelissetto & Vicari, *Phys Rep* 2002): $\delta = 4.789(2)$
- ▶ branch-point singularity at $\rho = \rho_c$.

critical isotherm approximant

We start with an approximant form that matches scaling behavior:

$$P_A = P_c \ll \sum_{n=0}^N A_n \rho^n \left(1 \ll \frac{\rho}{\rho_c}\right)^\delta$$

and follow steps from before to find the A_n 's:

1.

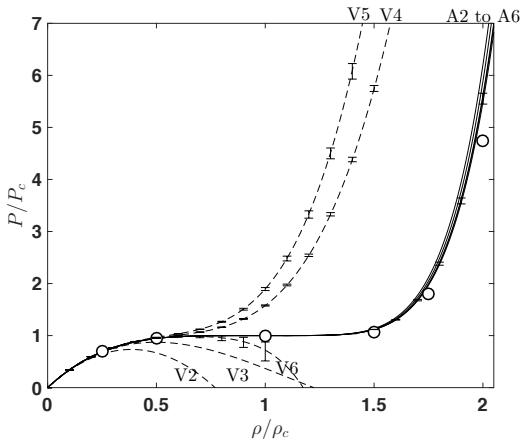
$$\left(P_c \ll kT_c \sum_{n=1}^N B_n(T_c) \rho^n\right) \left(1 \ll \frac{\rho}{\rho_c}\right)^{-\delta} = \sum_{n=0}^N A_n \rho^n$$

2. Expand LHS about $\rho=0$. JCP Miller's formula and Cauchy's product rule leads to a recursive solution:

$$A_n = \frac{P_c \Gamma(\delta + n)}{n! \rho_c^n \Gamma(\delta)} \ll \frac{kT_c}{\Gamma(\delta)} \sum_{j=0}^{n-1} \frac{B_{n-j} \Gamma(\delta + j)}{\rho_c^j j!}$$

critical isotherm approximant

$$P_A = P_c - \sum_{n=0}^N A_n \rho^n \left(1 - \frac{\rho}{\rho_c}\right)^\delta$$



(Barlow et al, JCP 2015)

Predict stuff from the approximant!

We have a critical isotherm approximant

$$P_A = P_c \ll \sum_{n=0}^N A_n \rho^n \left(1 \ll \frac{\rho}{\rho_c} \right)^\delta$$

with coefficients

$$A_n = \frac{P_c \Gamma(\delta + n)}{n! \rho_c^n \Gamma(\delta)} \ll \frac{kT_c}{\Gamma(\delta)} \sum_{j=0}^{n-1} \frac{B_{n-j} \Gamma(\delta + j)}{\rho_c^j j!}$$

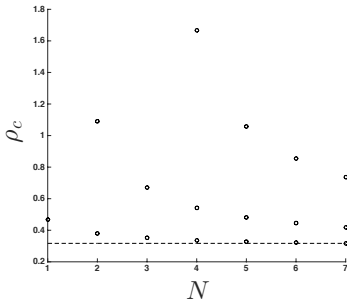
Recall that this could be posed as a system with N equations and N unknowns (A_0, A_1, \dots, A_N). Lets swap an unknown. Let one of the inputs (ρ_c, P_c, δ) be an unknown, keep N equations, and let the series have one less coefficient. This is equivalent to letting $A_N=0$ in above, leading to

$$P_c \ll \frac{kT_c N!}{\Gamma(\delta + N)} \sum_{j=1}^N \frac{\Gamma(\delta + N \ll j)}{(N \ll j)!} B_j \rho_c^j = 0$$

The above can be used to predict $P_c, \rho_c,$ or δ .

Predict stuff from the approximant!

$$P_c \ll \frac{kT_c N!}{\Gamma(\delta + N)} \sum_{j=1}^N \frac{\Gamma(\delta + N \ll j)}{(N \ll j)!} B_j \rho_c^j = 0$$



Given all other variables as inputs, the \circ 's are the fastest converging roots of the remaining ρ_c polynomial. The dashed line is the prediction from molecular simulations of a Lennard-Jones fluid.

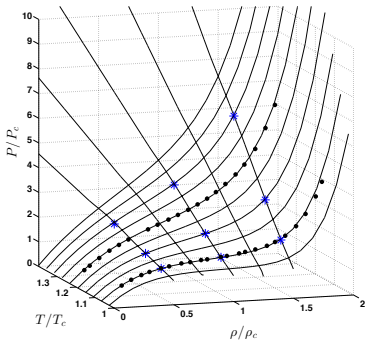
scaling → thermodynamic surface near c.p.

- ▶ along critical isotherm ($T = T_c$):

$$(P/P_c - 1) \sim \pm D_0 (1 - \rho/\rho_c)^\delta$$

- ▶ along critical isochore ($\rho = \rho_c$):

$$\left(\frac{\partial P}{\partial \rho}\right)_{T > T_c} \sim 1/\Gamma^+ (T/T_c - 1)^\gamma$$

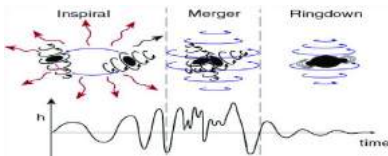


(Pelissetto & Vicari, 2002): $\delta = 4.789(2)$, $\gamma = 1.2372(5)$

(Barlow et al, JCP 2015)

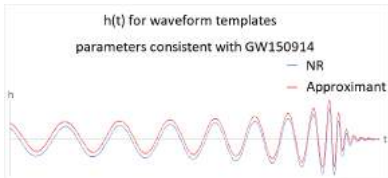
What's next on our plate?

Pre-2005 schematic:



K. S. Thorne, "Spacetime Warps and the Quantum World: Speculations about the Future," in R. H. Price, ed., *The Future of Spacetime*. W. W. Norton, New York, 2002, pp. 109-152.

Now thanks to Numerical Relativity:

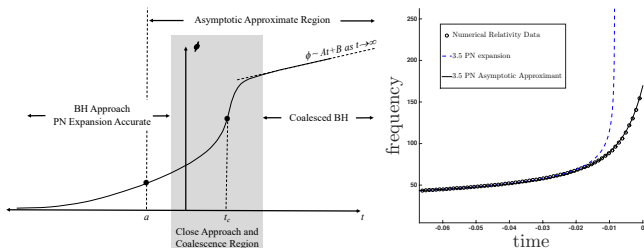


Taken from the Gravitational Wave Open Science Center

Gravitational Waveform Approximant

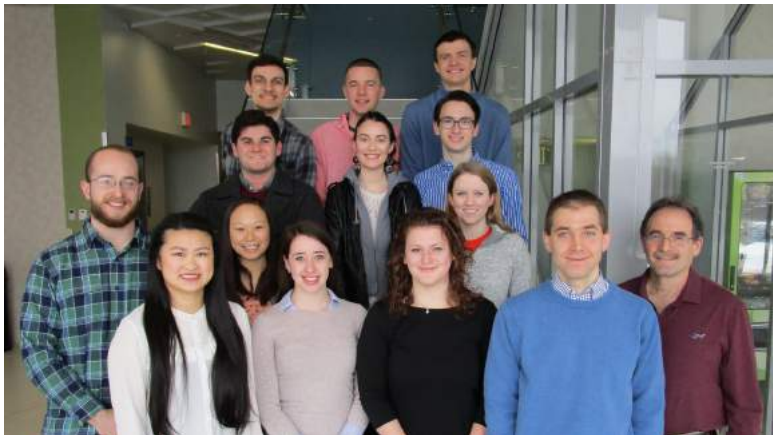
Preliminary Work:

- ▶ Expand known PN expansion about some negative time $t = a$
- ▶ Ultimate Goal: Form approximant that matches this new expansion at $t = a$ and matches known behavior as $t \rightarrow \infty$, and hopefully (as in problems of the past) pick up everything in-between! (cartoon on left)
- ▶ Where we are now: constructed approximant that matches expansion at $t = a$ and inflection at $t = 0$. (figure on right)
- ▶ Joint NSF/BSF proposal submitted with Ofek Birnholtz ... wish us luck!



On right: BHs of equal mass and zero spin. NR data in-house from Jam Sadiq.

Asymptotic Approximants Group



www.rit.edu/bwgroup

Asymptotic Approximants Group: References

Application to Astrophysics:

Beachley, Ryne J, Morgan Mistysyn, Joshua A Faber, Steven J Weinstein, and Nathaniel S Barlow. 2018. "Accurate Closed-Form Trajectories Of Light Around A Kerr Black Hole Using Asymptotic Approximants". Class. Quantum Grav. 35 (20)

Barlow, Nathaniel S, Steven J Weinstein, and Joshua A Faber. 2017. "An Asymptotically Consistent Approximant For The Equatorial Bending Angle Of Light Due To Kerr Black Holes". Class. Quantum Grav. 34 (135017)

Application to Thermodynamics:

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Application to Fluid Dynamics:

E. R. Belden, Z. A. Dickman, S. J. Weinstein, A. D. Archibee, E. Burroughs, and N. S. Barlow. 2020. "Asymptotic Approximant for the Falkner-Skan Boundary-Layer equation". Quarterly Journal Of Mechanics And Applied Mathematics. Accepted arxiv.org/abs/1907.09912

Barlow, Nathaniel S, Christopher R Stanton, Nicole Hill, Steven J Weinstein, and Allyssa G Cio. 2017. "On The Summation Of Divergent, Truncated, And Underspecified Power Series Via Asymptotic Approximants?". Quarterly Journal Of Mechanics And Applied Mathematics 70 (1): 21-48.